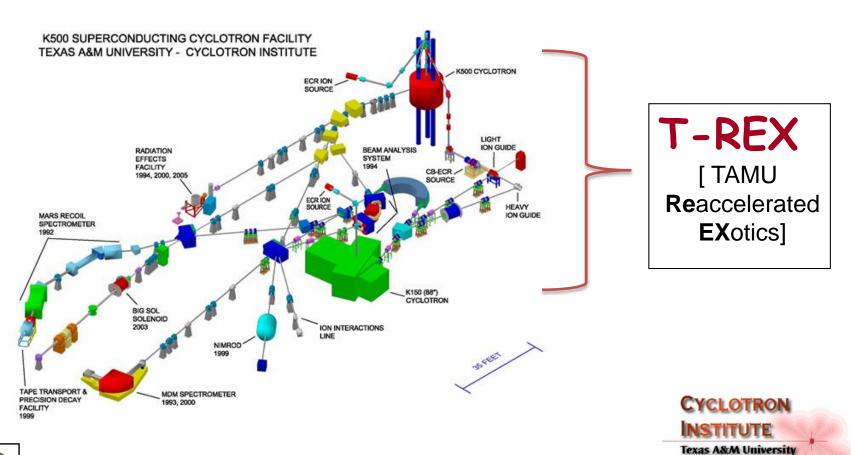
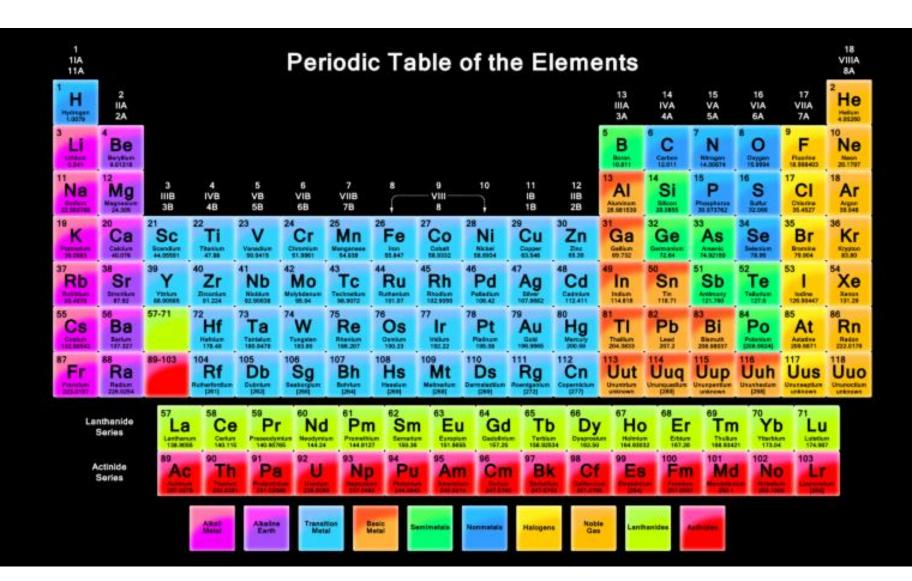
Nuclear Reactions Where Would We Be Without Them?

J. Natowitz



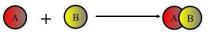






Types of Chemical Reactions

A) Synthesis Reactions: two reactants combine to make a larger product



B) Decomposition Reactions: larger reactant breaks down to form 2 or more simple products

 $\overrightarrow{AB} \longrightarrow \overrightarrow{A} + \overrightarrow{B}$

C) Single Displacement Reactions:

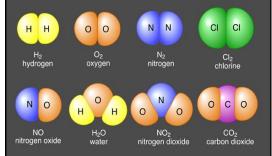
one element displaces or replaces an element in a compound

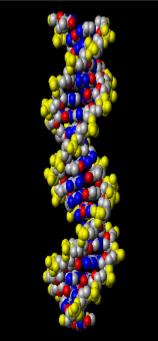


D) Double Displacement Reactions:

two elements in different compounds trade places

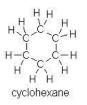




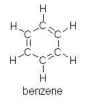


Some Organic Molecules H H-C HH H - H ННН methane ethane propane Н Н H ĤЙ ĤЙ Ĥ H H H butane pentane

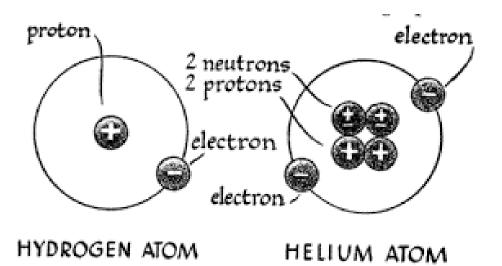
H H H H hexane



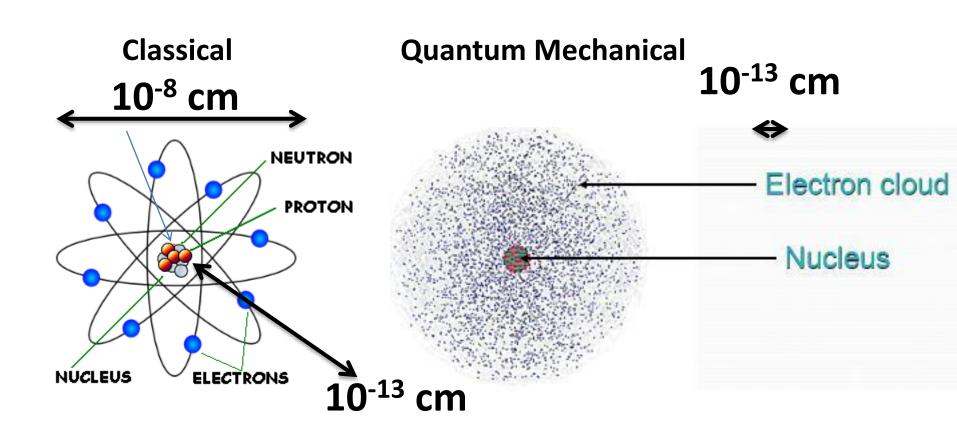
H H

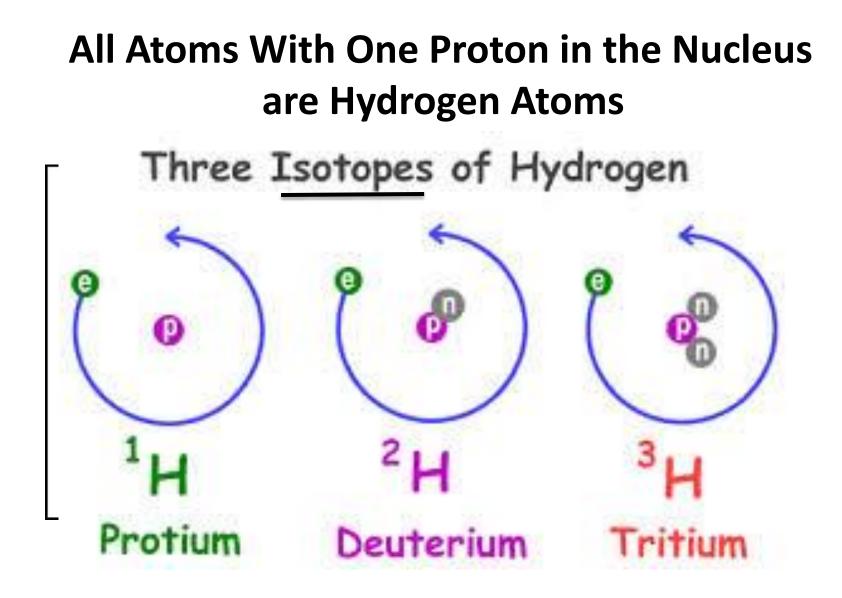


pentene

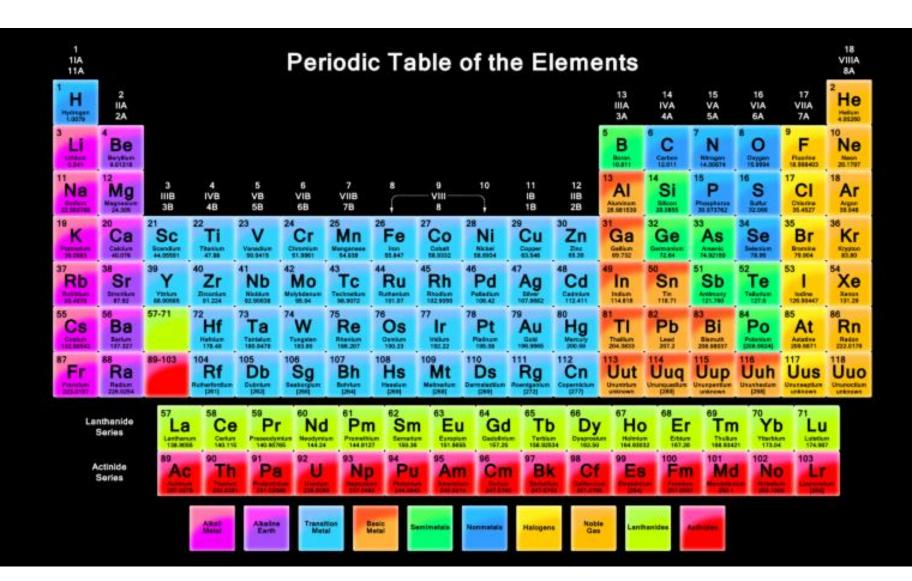


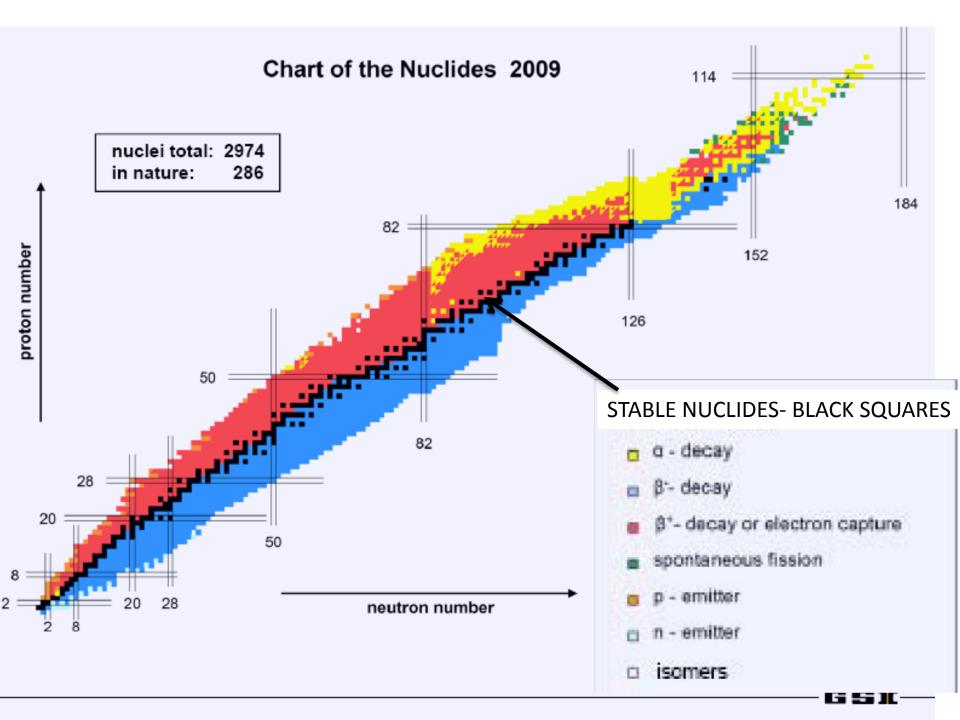
The Atom





There are many possible isotopes of each element.





Nuclear Binding Energy

Consider the formation of an atom from its constituents

$$Z \times {}^{1}H + N \times {}^{1}n \rightarrow {}^{A}X_{z}$$

 Define binding energy as the mass difference between the sum of the mass of free constituents (protons, neutrons, electrons) and the mass of the nucleus

$$E_B = [(Z \times M_H + N \times M_N - M(A,Z)]c^2]$$

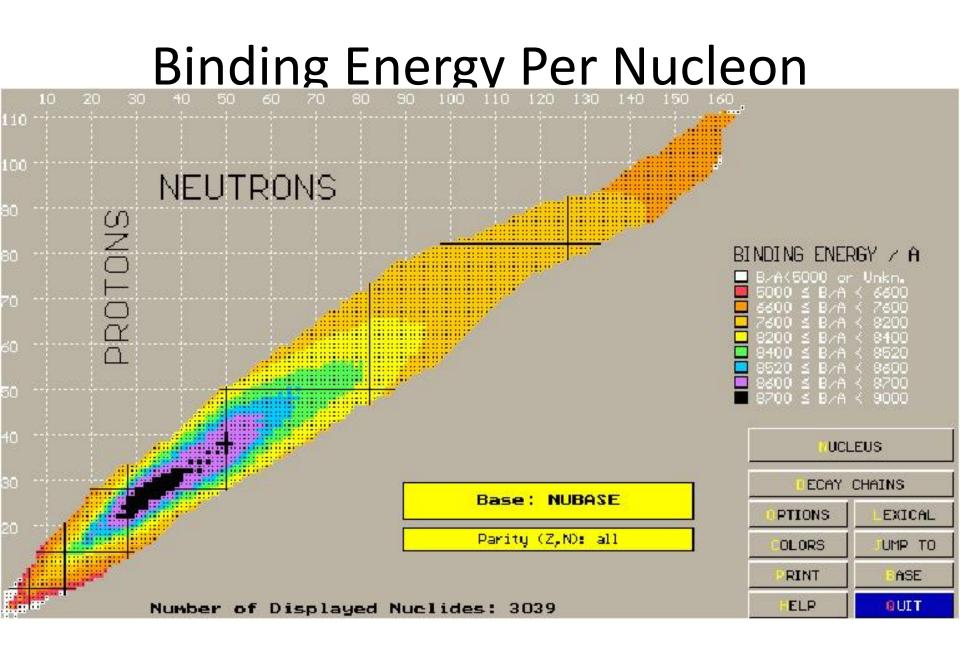
 $C^2 = 931.5 MeV/u$

Binding Energy

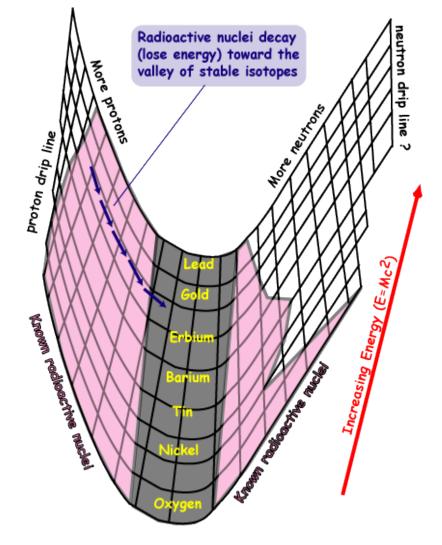
- The difference in mass between the constituent protons, electrons and neutrons and that of the isotope formed converted into energy, usually in MeV.
- For 208 Pb BE = 1636 MeV

Binding Energy per Nucleon

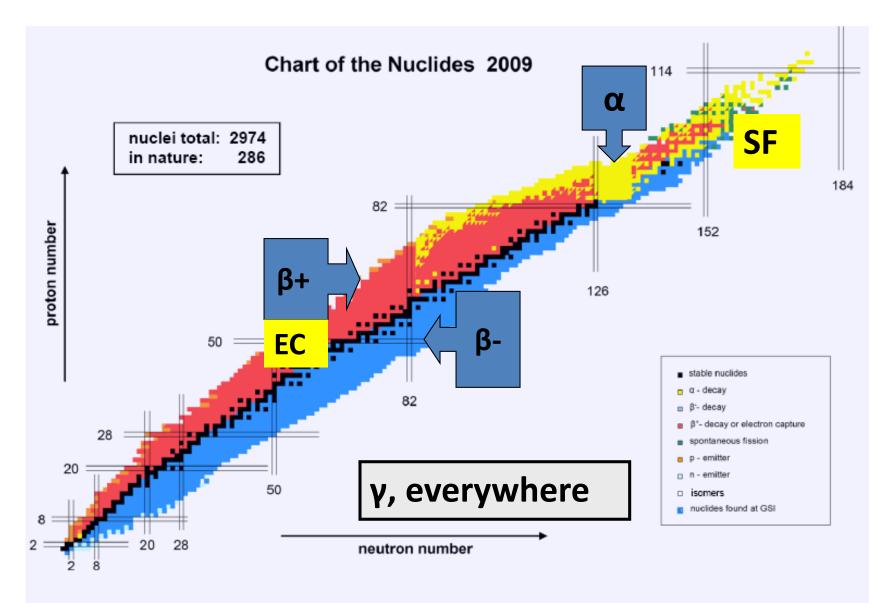
- The Binding energy divided by A, the total number of nucleons.
- For ²⁰⁸Pb BE = 1636/208 = 7.87 MeV/nucleon



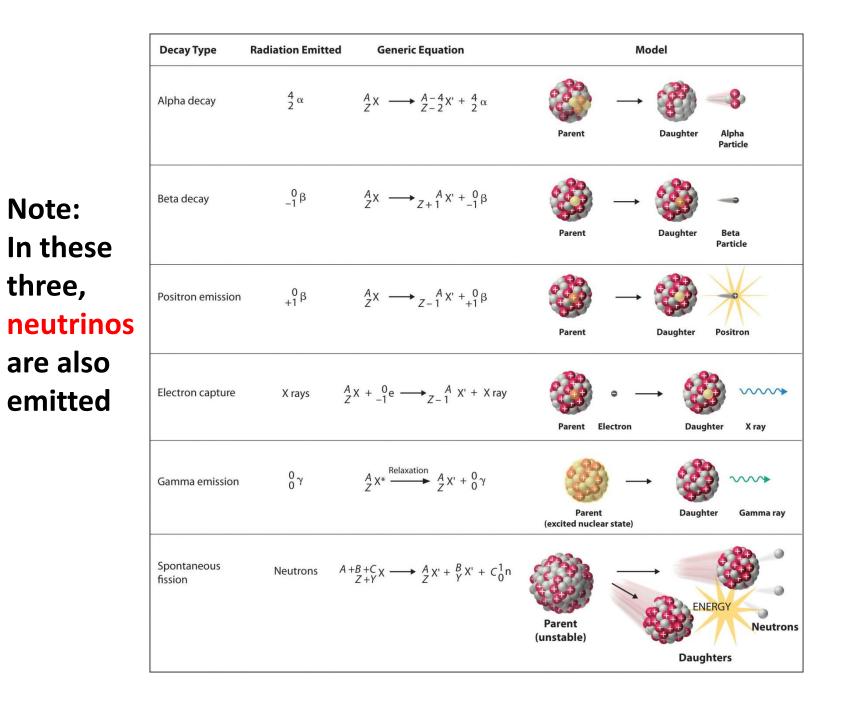


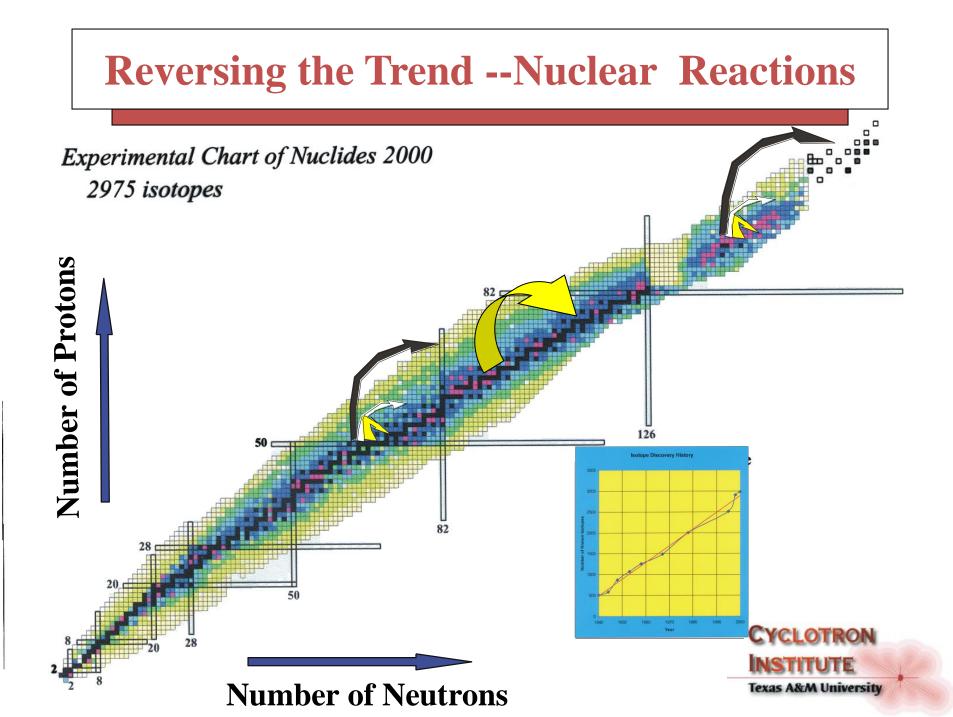


Most Isotopes are Radioactive

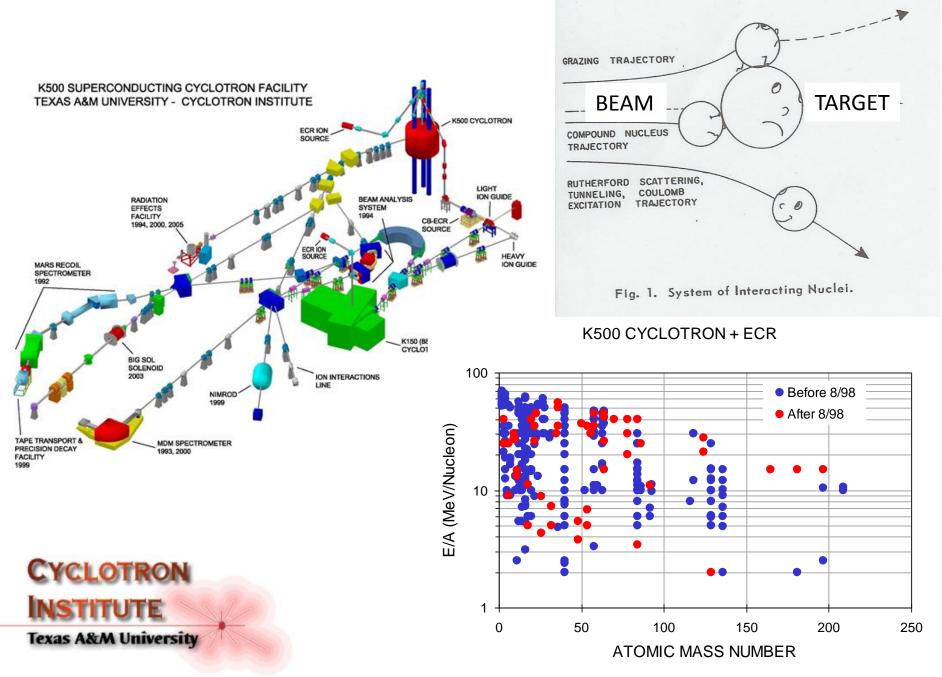


GSI



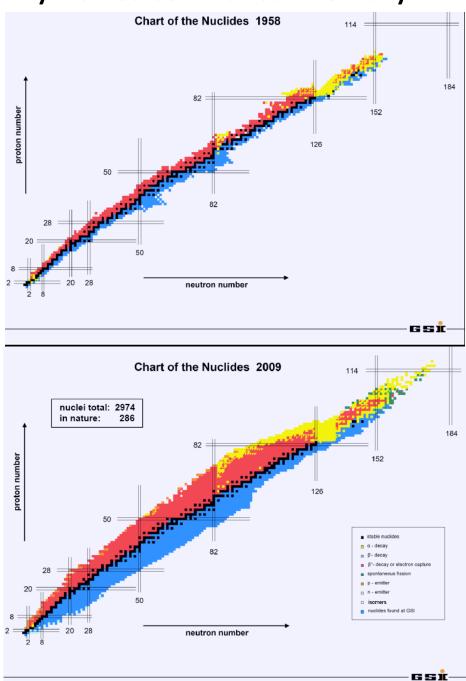


NEED AN ACCELERATOR

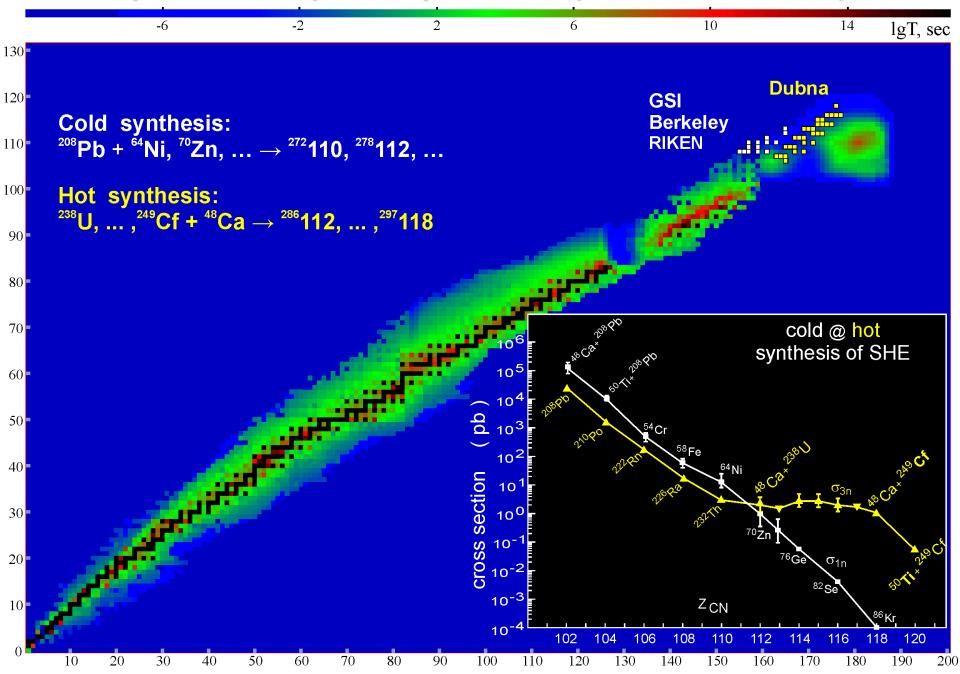


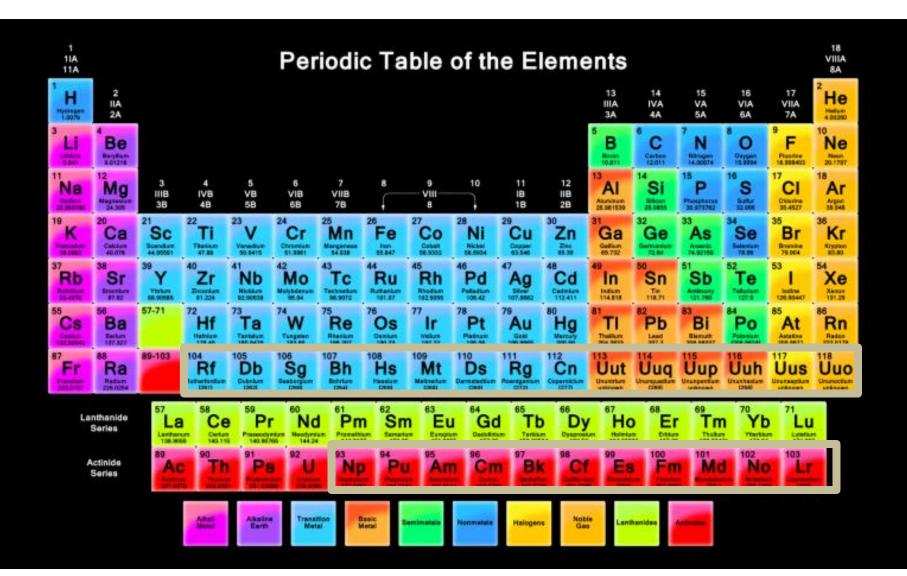
1958-My first course in nuclear chemistry

 With such reactions new isotopes
 and
 new elements
 can be made

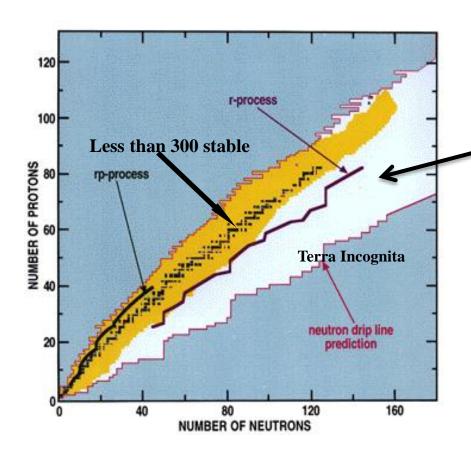


Synthesis of superheavy elements (cold and hot fusion)

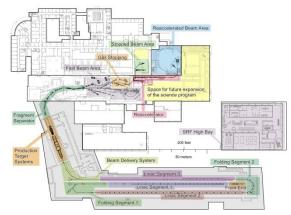




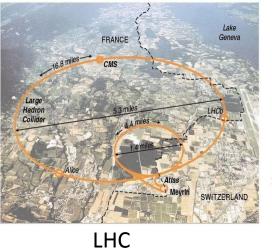
Theoretical Limits to the Existence of Nuclei



- Only a fraction of the theoretically possible isotopes have been produced and studied.
- A new generation of accelerators
 being constructed will accelerate radioactive ions and probe the region of unknown isotopes



Layout of the accelerator, experimental systems and the experimental areas of the Facility for Rare Isotope Beams.

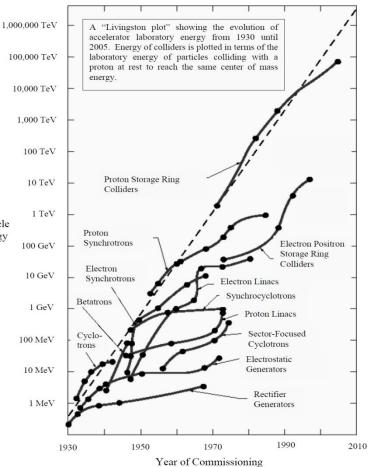




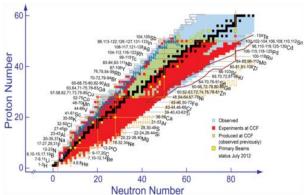
RHIC



ACCELERATORS

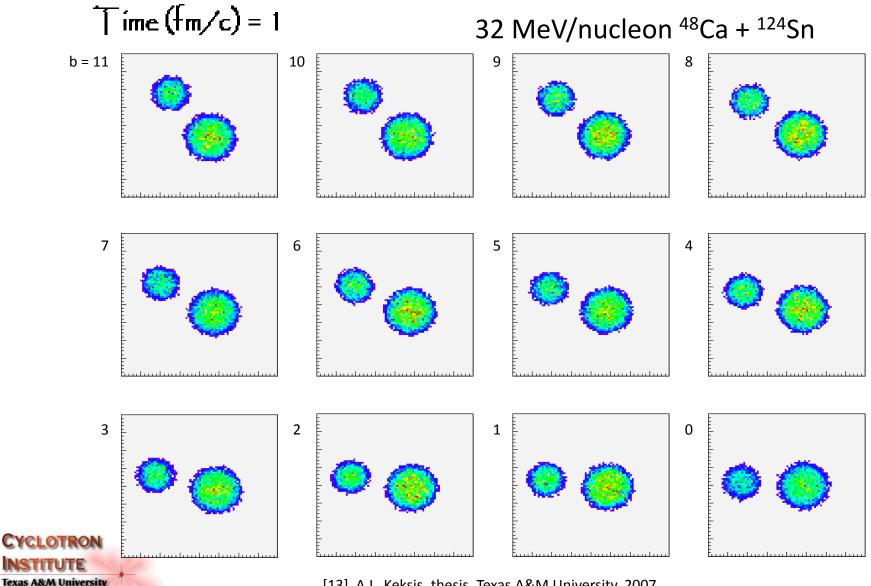




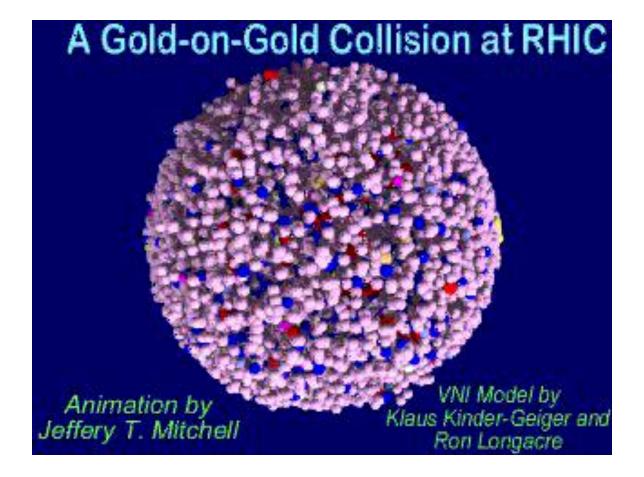


TAMU

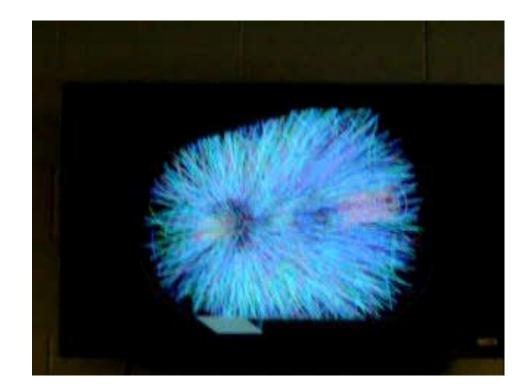
Higher Energy Reactions



[13] A.L. Keksis, thesis, Texas A&M University, 2007



The Little Bang

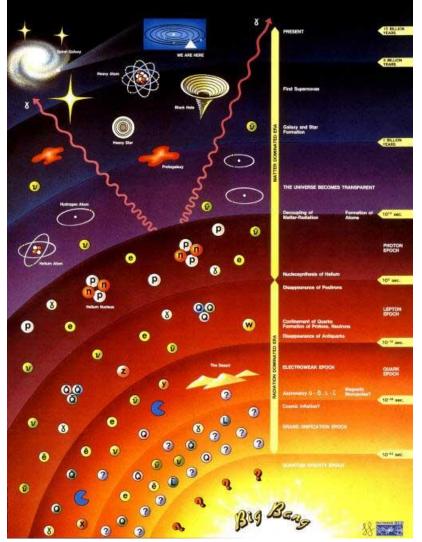


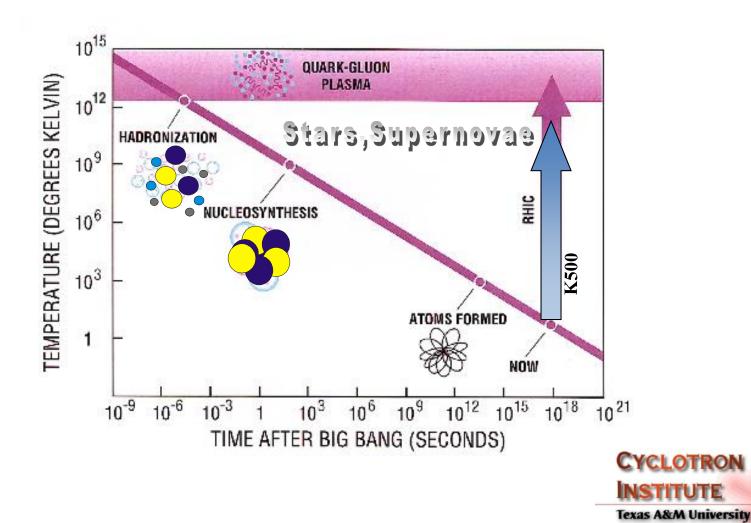


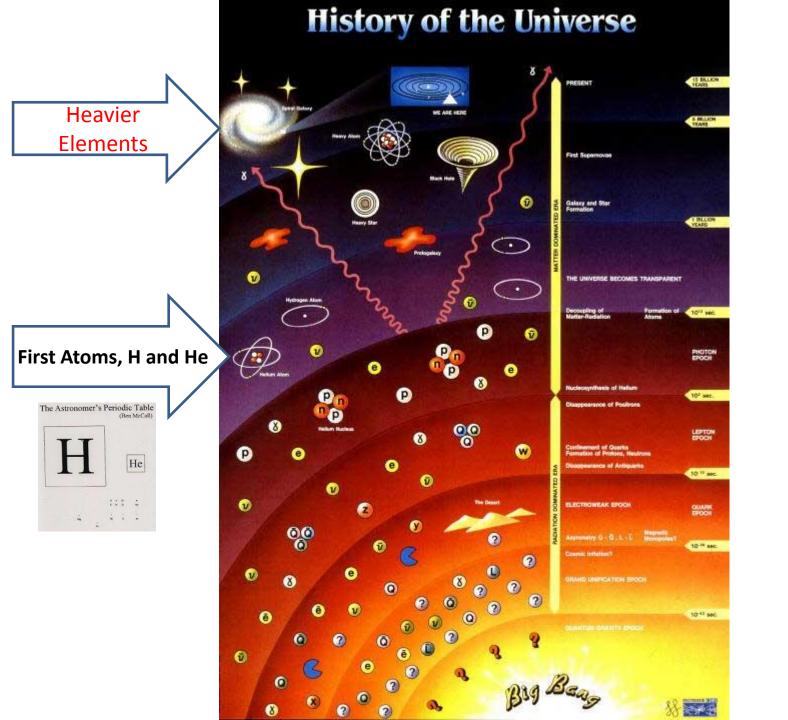




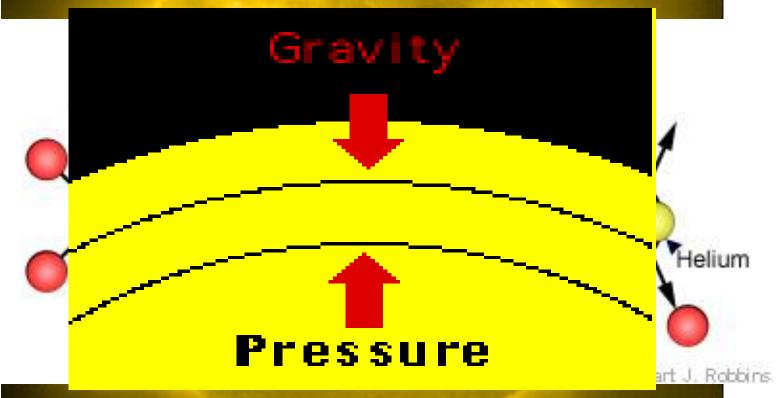
History of the Universe



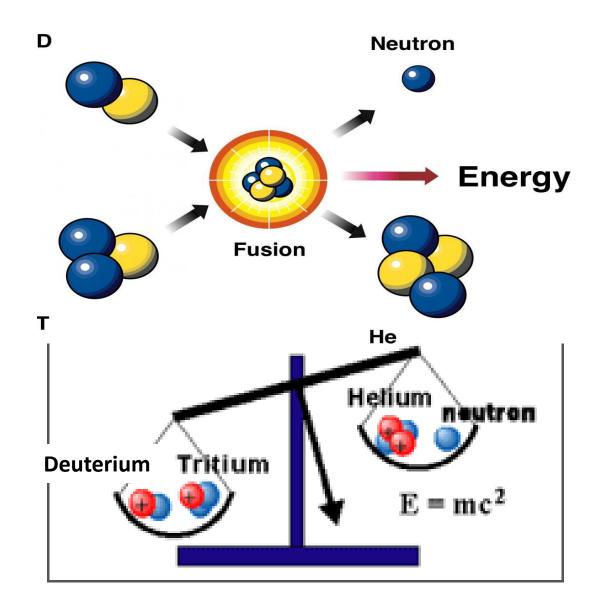




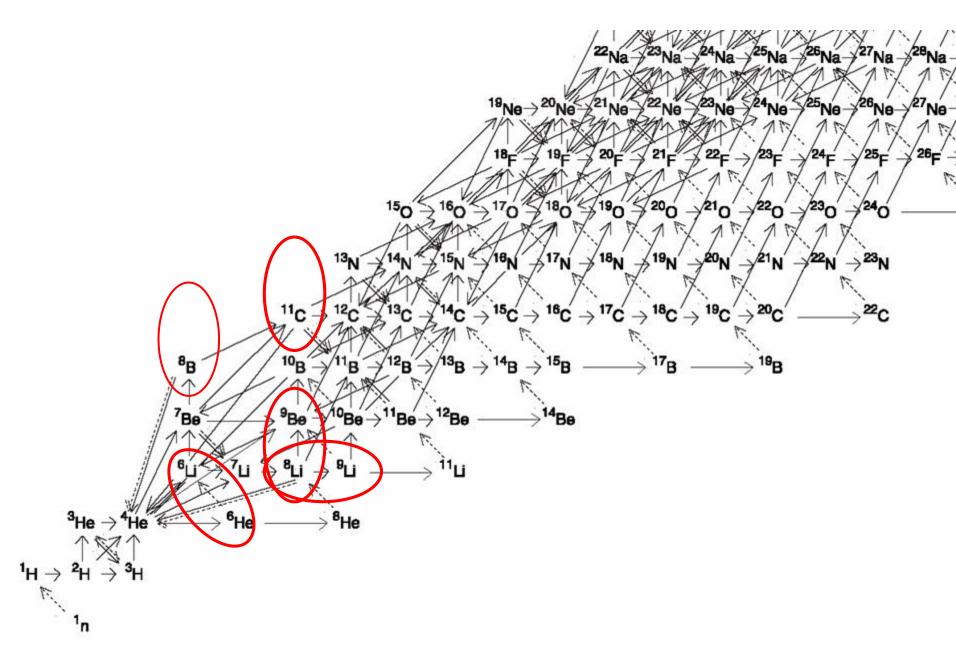
OUR SUN A Ball of Hydrogen



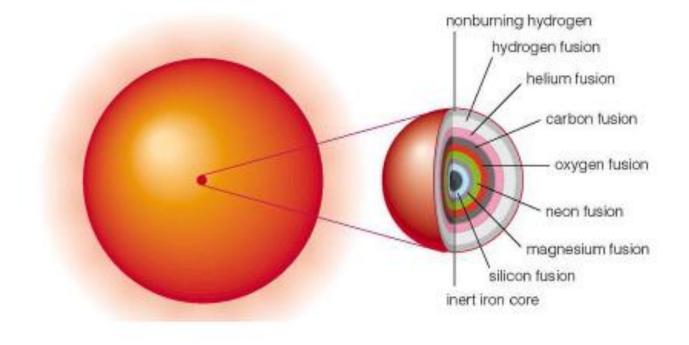
10,000,000 DEGREES KELVIN 2001/03/04 01:05



Origin of the Lighter Elements

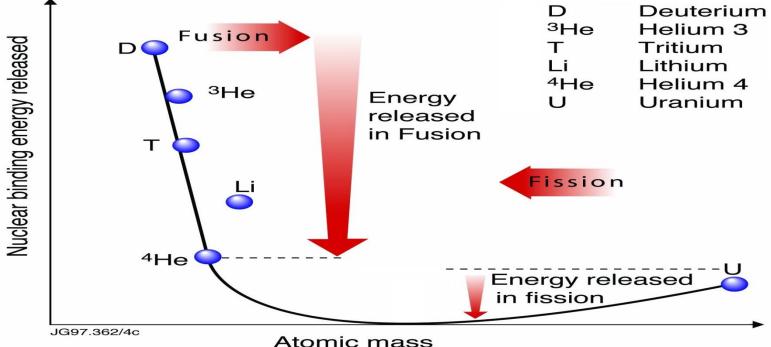


A massive star near the end of its lifetime has "onion ring" structure

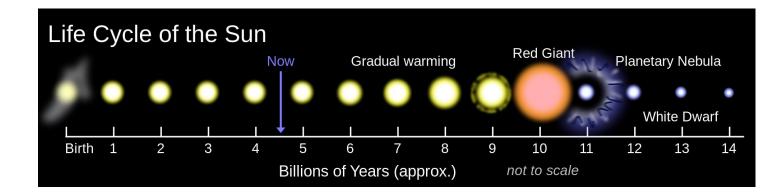


WHY? Because Floor of Binding Energy Valley is Not Flat

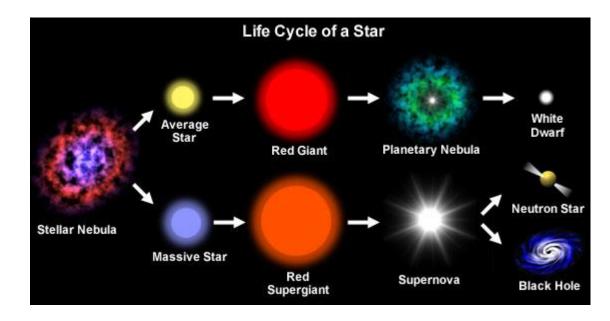
For A Given Number of Nucleons, An Increase in Binding Energy per Nucleon Means that Energy Will Be Released

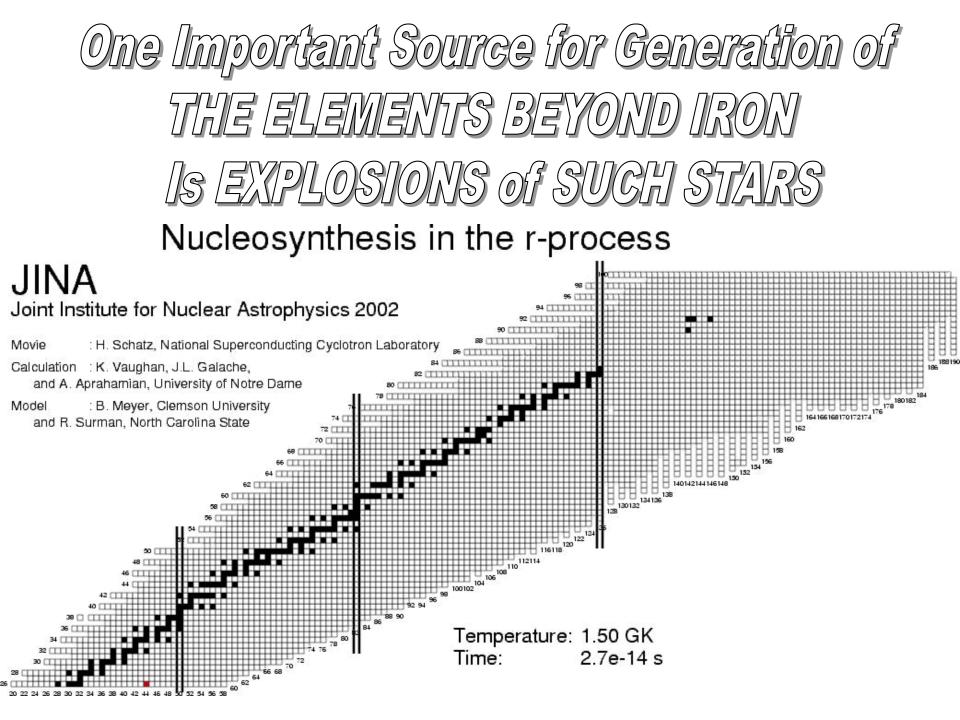


For Small Numbers of Nucleons, i.e., Nuclei Much Lighter Than Iron, Energy is Released If Nuclei Fuse (FUSION) For Large Numbers of Nucleons i.e., Nuclei Much Heavier Than Iron, Energy is Released If the Nuclei Split (FISSION)



HOW ARE THE HEAVIER ELEMENTS PRODUCED ?

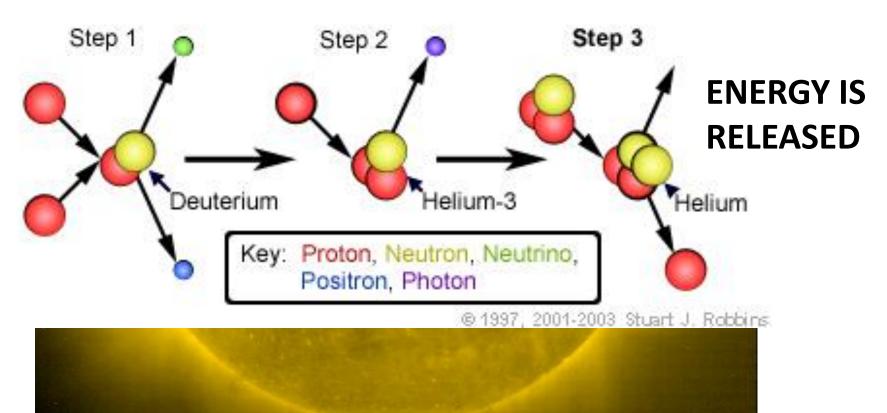




H B	Big Large Super- Bang Lars novae														
Li Be	Cosmic Small Man- rays stars made										C S L	N s L	O S L	F	Ne s L
											Si s L	P L	S S L	CI	Ar
K Ca Sc	Ti s L	V \$ L	Cr	Mn	Fe ^{\$ L}	Co \$	Ni \$	Cu	Zn	s L Ga	Ge	As	Se \$	Br \$	Kr \$
Rb Sr Y	Zr	Nb	Mo \$ L	Tc L	Ru s L	Rh \$	Pd s L	Ag \$ L	Cd \$ L	In \$ L	Sn \$ L	Sb \$	Te \$	\$	Xe \$
Cs Ba	Hf \$ L		W \$ L	Re \$	Os \$	lr \$	Pt \$		Hg \$ L	TI \$ L	Pb \$	Bi \$	Po \$	At \$	
Fr Ra	La Ce Pr Nd Pm Sm Eu Gd Tb Dy Ho Er Tm Yb Lu														
		L Th	s L Pa	\$ L	s L Np	\$ L	\$	s Cm	\$	\$	\$	s Fm	\$	\$ L No	\$
	\$	\$	\$	\$	\$	\$	М	M	M	M	M	м	M	м	M

OUR SUN

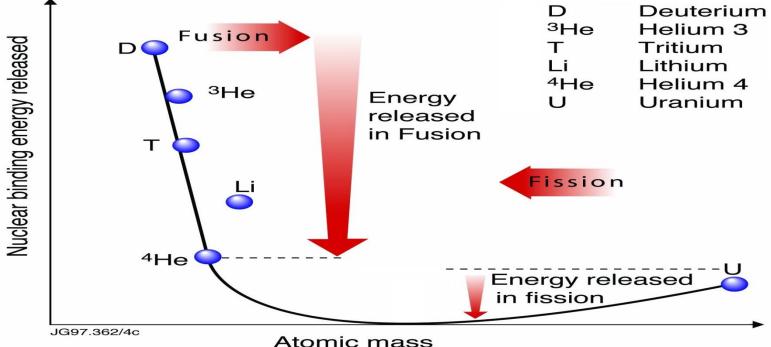
Nuclear Fusion in Stars



10,000,000 DEGREES KELVIN 2001/03/04 01:05

WHY? Because Floor of Binding Energy Valley is Not Flat

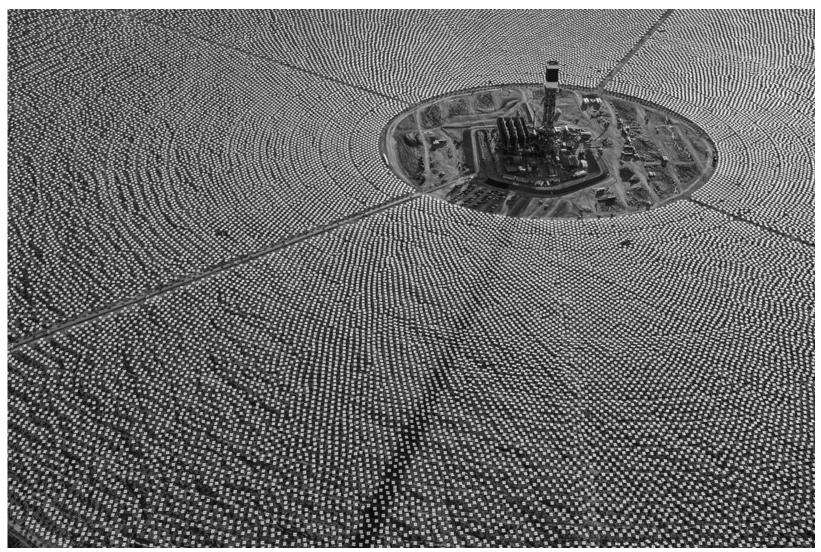
For A Given Number of Nucleons, An Increase in Binding Energy per Nucleon Means that Energy Will Be Released



For Small Numbers of Nucleons, i.e., Nuclei Much Lighter Than Iron, Energy is Released If Nuclei Fuse (FUSION) For Large Numbers of Nucleons i.e., Nuclei Much Heavier Than Iron, Energy is Released If the Nuclei Split (FISSION)

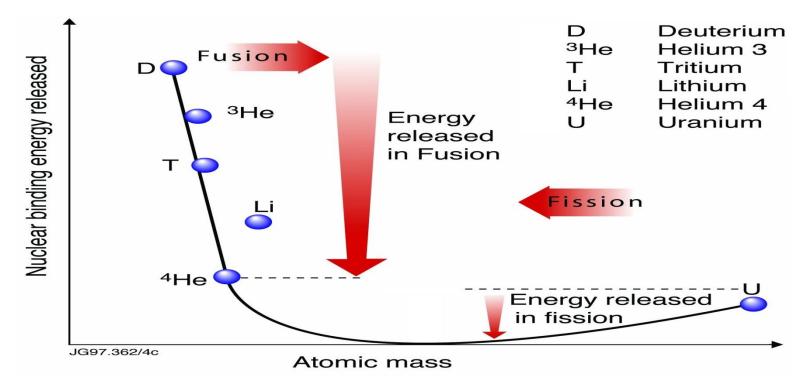
Using Fossil Fuels is Indirectly Harvesting Solar Energy For Energy Production

Directly Harvesting Solar Energy For Energy Production



Harnessing Nuclear Reactions For Energy Production

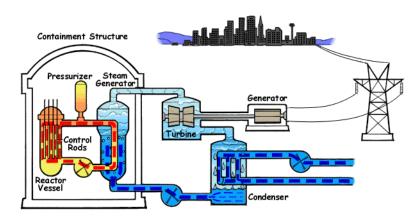
For A Given Number of Nucleons, An Increase in Binding Energy per Nucleon Means that Energy Will Be Released

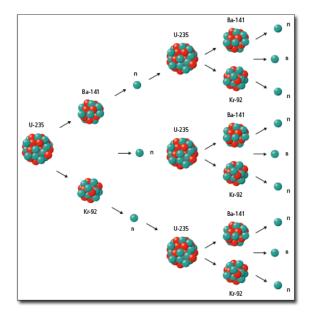


For Small Numbers of Nucleons, i.e., Nuclei Much Lighter Than Iron, Energy is Released If Nuclei Fuse (FUSION) For Large Numbers of Nucleons i.e., Nuclei Much Heavier Than Iron, Energy is Released If the Nuclei Split (FISSION)

FISSION REACTOR Harnessing the Heavy Elements

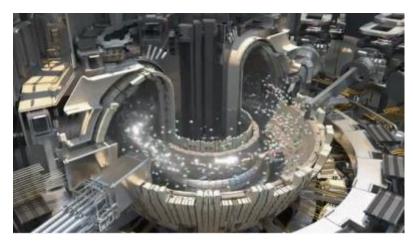
NOW

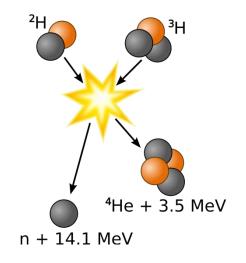




FUSION REACTOR Harnessing the Light Elements

Still in the future





Nuclear Reactions Where Would We Be Without Them?



"There are more things in heaven and earth, Horatio, Than are dreamt of in your philosophy." Shakespeare

- Hamlet (1.5.167-8), Hamlet to Horatio